

Canada's national laboratory for particle and nuclear physics and accelerator-based science

# Decay Spectroscopy of Neutron-Rich Cadmium Around the N = 82 Shell Closure

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IX Tastes of Nuclear Physics, UniZulu

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What are atoms made of?

How were the elements made?



#### What are atoms made of?

1949: Shell model of nuclear structure

How were the elements made?

- 1939-1948: Big Bang and Stellar nucleosynthesis
- 1957: Heavy element nucleosynthesis

Today: study of exotic radioactive nuclei at accelerator facilities!



Increased binding energy at particular proton (Z) and neutron (N) numbers



- Full shell (closures) occur at magic numbers
- Closed proton shell and closed neutron shell in **doubly-magic** nucleus



Nucleon-nucleon interaction and mean-field potential

$$H = \left[\sum_{i=1}^{A} \frac{p_i^2}{2m} + \sum_{i=1}^{A} v(\vec{r_i})\right] + \left[\sum_{i \neq k}^{A} V(\vec{r_{i,k}}) - \sum_{i=1}^{A} v(\vec{r_i})\right]$$
$$= H^0 + W_{RES}$$



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W<sub>RES</sub> = 0: Non-interacting shell model

$$H^{0} = \sum_{i=1}^{A} \frac{p_{i}^{2}}{2m} + \sum_{i}^{A} v\left(\vec{r_{i}}\right)$$

Choice of potential gives single-particle energy levels

$$V(r_i) = \frac{-V_0}{1 + \exp\left[(r_i - R)/a\right]} + V_{so}(r_i) \ \vec{l} \cdot \vec{s}$$

Some ground state properties (J<sup>π</sup>) are reproduced





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$$= H^0 + W_{RES}$$

- *W<sub>RES</sub>* ≠ 0: Interacting shell model
  - Configuration mixing
  - Effective two-body (NN) matrix elements
- Recent developments:
  - Chiral effective field theory (3N)
  - Ab initio methods (many-body)



Valence space



- pp-cycle and CNO-cycle up to <sup>16</sup>O, advanced burning up to <sup>56</sup>Fe
- Waiting point nuclei form the rapid neutron capture (*r*-) process path, which create half of the nuclei heavier than <sup>56</sup>Fe





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- Waiting point nuclei form the rapid neutron capture (*r*-) process path, which create half of the nuclei heavier than <sup>56</sup>Fe
- N = 82 isotope <sup>130</sup>Cd provides critical information on the **abundance peak** at  $A \sim 130$ .



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Supernova explosion (hot and cold *r*-process winds), and neutron star merger

$$\frac{n\left(A+1,Z\right)}{n\left(A,Z\right)} = n_n \cdot \sqrt[3]{\frac{A+1}{A}} \cdot \frac{2\pi\hbar^2}{k_B T \cdot m_u} \cdot \exp\left(-\frac{S_n}{k_B T}\right)$$





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- Defining nuclear physics parameters
  - Masses (S<sub>n</sub>, Q<sub>β</sub>)
  - β-decay lifetimes
  - β-delayed neutron emission probabilities rates
  - Shell structure far off stability



*r*-process sites



## Neutron-rich nuclear decay





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### Neutron-rich nuclear decay



Selection rules of  $\beta$ -decay

			Fermi	Gamow-Teller	
Transition Type	L	$\Delta \Pi$	$\Delta I$	$\Delta I$	$\log(ft)$
Allowed	0	No	0	(0), 1	$\sim 4.0 - 7.5$
First Forbidden	1	Yes	(0), 1	$0,\!1,\!2$	$\sim 6.0 - 9.0$
Second Forbidden	2	No	(1), 2	$^{2,3}$	$\sim 10 - 13$



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Transition Type

Allowed

First Forbidden

Second Forbidden

 $\Delta \Pi$ 

No

Yes

No

L

0

 $\mathbf{2}$ 

 $\Delta I$ 

0

(0),1

(1),2

 $\Delta I$ 

(0),1

0,1,2

2,3

#### Neutron-rich nuclear decay



 $\log(ft)$ 

 $\sim 4.0 - 7.5$ 

 $\sim 6.0 - 9.0$ 

 $\sim 10 - 13$ 

 $|J_i - J_f| \leqslant L \leqslant J_i + J_f$  $\Pi(EL) = (-1)^L$  $\Pi(ML) = (-1)^{(L+1)}$ 



 <sup>128-132</sup>Cd are neighboring the doubly-magic <sup>132</sup>Sn, which is central to shell model calculations and *r*-process simulations

	1288n 59.07 M	1298n 2.23 M	130Sn 3.72 M	131Sn 56.0 S	132Sn 39.7 S	133Sn 1.46 S	134Sn 1.050 S	1358n 530 MS	136Sn 0.25 S	
Z	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00% β-n: 0.03%	β-: 100.00% β-n: 17.00%	β-: 100.00% β-n: 21.00%	β-: 100.00% β-n: 30.00%	
49	127In 1.09 S	128In 0.84 S	129In 0.61 S	130In 0.29 S	131In 0.28 S	132In 0.207 S	133In 165 MS	134In 140 MS	135In 92 MS	
	0 . 100 0004		570 (10) ms	284 (10) ms	261 (3) ms	198 (2) ms	163 (7) ms	126 (7) ms	103 (5) ms	
	β-n≤ 0.03%	810 (30) ms	β-n: 0.25%	β-n: 0.93%	β-n≤ 2.00%	β-n: 6.30%	β-n: 85.00%	β-n: 65.00%	β-n	
48	126Cd	127Cd	128Cd	129Cd	130Cd	131Cd	132Cd	133Cd		
	0.515 \$	0.37 5	0.28 \$	0.27 \$	162 MS	68 MS	97 MS	57 MS		
	β-: 100.00%	β-: 100.00%	245 (5) ms	154 (8) ms 151 (15) ms	127 (2) ms β-n: 3.50%	98 (2) ms β-n: 3.50%	82 (4) ms β-n: 60.00%	64 (8) ms β-n		
	125Ag	126Ag	127Ag	128Ag	129Ag	130Ag				
	166 MS	107 MS	109 MS	58 MS	46 MS	≈ 50 MS				
47	β-: 100.00% β-n	β-: 100.00% β-n	β-: 100.00%	β-: 100.00% β-n	β-: 100.00% β-n	β-n β-				
46	124Pd 38 MS	125Pd >230 NS	126Pd >230 NS		128Pd >394 NS	Recent t <sub>1/2</sub> measurements:				
	β-: 100.00%	β-n β-	β-n β-		β-n β-	Lorusso et al., PRL 114, 192501 (2015) Taprogge et al., PRC 91, 054324 (2015)				
	78	79	80	81	82	83	84	85	N	



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	126Cd	127 Cd	128Cd	1.25 Cd	130C.	131Cd	132Cd	1 33Cd	
40	0.515.5	0. <i>r</i> 1	0.26 5	.2. 5	102.15	08 (2) mc	97 MS	5 145	
48	β-: 100.00%	β-: 100. <b>0%</b>	245 (5) ms	15   (8   ms 15   (7 5) ms	β-n: 3. 10%	β-n: 3.50%	82 (4) ms β-h: 60.00%	64 (8) ms β-h	
	125Ag 166 MS	126A 107 MS	127Ag	1 28Ag 58 MS	129Ag 46 MS	⇒ 50 ma			
47	β-: 100.00% β-n	β-: 100.00% β-n	β 100.00%	β-: 100.00% β-n	β-: 100.00% β-n	β-n β-			
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	78	79	80	81	82	83	84	85	N



- 1988 experiment in Sweden: 7 transitions and 4 levels (1 isomer)
- Re-evaluated in 2015.



Work by B. Ekstrom quoted in B. Fogelberg, Nucl. Data for Sc. and Tech., **837** (1988)

Z. Elekes and J. Timar, Nucl. Data Sheets 129, 191 (2015)



- Selective ionization with the Ion Guide Laser Ion Source [IG-LIS]
  - Background suppression by factors 10<sup>5</sup>-10<sup>6</sup>







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  - Background suppression by factors 10<sup>5</sup>-10<sup>6</sup>
- High statistics β-γ-γ with SCEPTAR : SCintillating Electron Positron Tagging Array







In-vacuum moving tape collector system



- Selective ionization with the Ion Guide Laser Ion Source [IG-LIS]
  - Background suppression by factors 10<sup>5</sup>-10<sup>6</sup>
- High statistics β-γ-γ with SCEPTAR : SCintillating Electron Positron Tagging Array
- 16 large-volume germanium GRIFFIN detectors dedicated to decay spectroscopy of the low-energy radioactive ion beams at TRIUMF.







In-vacuum moving tape collector system





Further discrimination of isobaric background:

 Identification of transitions by comparing laser on (Cd + In) and laser blocked (mostly In)



β-gated γ-singles

\* Known <sup>128</sup>Cd→<sup>128</sup>In

\* *New* <sup>128</sup>Cd→<sup>128</sup>In









### β-γ-γ coincidence <sup>12</sup>





32 new transitions and 11 new states





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- 32 new transitions and 11 new states
- Structure information from log(ft) values and angular correlations





- NuShellX (jj45pna)
- In-Medium Similarity Renormalization Group (IMSRG)





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- In-Medium Similarity Renormalization Group (IMSRG)





• 7 transitions observed in <sup>131</sup>In at ISOLDE, **23** at RIKEN: only **4** transitions in common



J. Taprogge et al., Eur. Phys. J. A 52, 347 (2016)



- 7 transitions observed in <sup>131</sup>In at ISOLDE, 23 at RIKEN: only 4 transitions in common
- No  $\gamma$ -transitions observed from the  $\beta$ -decay of <sup>132</sup>Cd to <sup>132</sup>In.



J. Taprogge *et al.*, Eur. Phys. J. A **52**, 347 (2016)



J. Taprogge et al., Phys. Rev. Lett. 112, 132501 (2014)



### $\beta$ -decay of <sup>131</sup>Cd <sup>16</sup>

- Delivered at 0.7 pps
- Many transitions confirmed:
   5/7 (O. Arndt *et al.*, Acta Phys. Pol. B, 2009)
   21/23 (J. Taprogge *et al.*, Eur. Phys. J, 2016)
- One state in <sup>130</sup>In strongly populated by βn-decay





- Only one coincidence relationship observed
- 4 transitions placed based on energy differences





- 1 proton away from double shell closure
- Only single-particle excited states can be calculated





- Very low neutron separation energy → large neutron branching ratio
  - 988 keV transition expected in both <sup>131-132</sup>Cd datasets





#### **Decay Spectroscopy** of Neutron-Rich Cadmium Around the N = 82 Shell Closure

- <sup>128</sup>In: 32 new transitions and 11 new levels
- <sup>131</sup>In: 8 transitions placed and 8 levels
- <sup>132</sup>Cd: Higher yields and cleaner beams required
- Would not have been possible without the discriminating power of IG-LIS
- Important inputs for theoretical models of the *r*-process and the nuclear structure of exotic isotopes
- Understanding of nuclear forces and shell evolution in a region which only recently became accessible for more extensive studies!



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# Thank you!

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